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# Amalytical solluticm of the Schrëdimger equation with linear confimememt potential 

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#### Abstract

It is shown that the Schrodinger equation wath linear confinement potential has an analytical solutson The existence of che analytical solution requires additional terms to the potentral other than the inear one. The Coulomb-type potental is necessary to get non-zero energy eigenvalues We propose an asymptoticaliy linear potential, and show that we can analytically colve the Schrodinger equation with this potential


## 1. Hefroduction

Quarks are believed to be confined to baryons or mesons. Many experimental data show that the confinement potential is proportional to the distance between quarks at large distance. There have been many attempts [1] to deduce confinement potentials diredly from the QCD Lagrangian, but no-one has succeeded.

The moturn of quarks in a harmonic oscillator type potential [2], proportional to the square of the distance or in the bag [3], has been well studied. On the other hand, hadr in masses and heavy quarkonia have been numerically discussed with linear comanement potentials [4-6]. However, an analytical solution of the quark equation of motion with an asymptotically linear potential proportional to the distance between quarks at large distance has not yet been derived

In this paper one analytical solution of the Schrodinger equation with a linear potential proportional to the distance is suggested. It is well known that the solution of the Schrödinger equation with a harmonic oscillator type potential proportional to the square of the distance $r^{2}$ can be written as the product of polynomial functions of $r$ and exp $\left(-r^{2}\right)$, and the one with a Couiomb type potential proportional to the inverse of the distance $1 / r$ is writen as the product of polynomial functions and $\exp (-r)$. The solution of the Schrödinger equation with a linear confinement potential is naturally supposed to have the form of $\exp \left(-r^{n}\right)$, where $1<n<2$. Actually it is shown that the solution with $n=\frac{3}{2}$ satisfies the Schrödinger equation with an asymptotically linear confinement potential. At the same time, the existence of an analytical solution requires additional terms to the usual confinement potential which is made up of linear + Coulomb terms.

The trial potential we use is the linear + Coulomb type potential. Not only is this potential very popular as a phenomenological potential [4], but also the Coulomb term is aecessary for the Schrödinger equation to have a solution with non-zero energy eigenvalues, as discussed later.

## 2. Deduction of recurreme formula

Spherical symmetric Schrödinger equation for a particle with mass $m$ and orbital angular momentum $l$ is given by

$$
\begin{equation*}
\frac{1}{r^{2}} \frac{\mathrm{~d}}{\mathrm{~d} r}\left(r^{2} \frac{\mathrm{~d} R}{\mathrm{~d} r}\right)-\frac{l(l+1)}{r^{2}} R+\frac{2 m}{\hbar^{2}}\{E-U(r)\} R=0 \tag{2.1}
\end{equation*}
$$

where $R(r)$ is the radial part of the wavefunction $\psi(r)$;

$$
\begin{equation*}
\psi(r)=R(r) Y_{l}^{m}(\Omega) . \tag{2.2}
\end{equation*}
$$

The energy eigenvalue $E$ of the particie is determmed by solving (21).
For the spherical symmetric potential $U(r)$, let us consider the confinement potential which is made of a long range linear potential and a short range Coulomb type potential [4]:

$$
\begin{equation*}
U(r)=a r-\frac{b}{r} \tag{2.3}
\end{equation*}
$$

Constants $a$ and $b$ are independent of distance $r$, and their values are determined by experiments so as to reproduce proper physical values of the hadrons. Then equation (2.1) is rewritten as

$$
\begin{equation*}
\frac{1}{r^{2}} \frac{\mathrm{~d}}{\mathrm{~d} r}\left(r^{2} \frac{\mathrm{~d} R}{\mathrm{~d} r}\right)=\left\{\frac{2 m}{\hbar^{2}} a r-\frac{2 m}{\hbar^{2}} E-\frac{2 m}{\hbar^{2}} \frac{b}{r}+\frac{l(l+1)}{r^{2}}\right\}< \tag{2.4}
\end{equation*}
$$

At large $r$, equation (2.4) becomes

$$
\begin{equation*}
\frac{\mathrm{d}^{2} u}{\mathrm{~d} r^{2}}=\frac{2 m}{\hbar^{2}} a r u-\frac{2 m}{\hbar^{2}} E u \tag{25}
\end{equation*}
$$

where

$$
\begin{equation*}
R(r)=\frac{u(r)}{r} \tag{2.6}
\end{equation*}
$$

If Wh requre the term proportional to $r$ after double differentiation, $u(r)$ must have the form $\exp \left(-r^{3 / 2}\right)$ and for the constant term, $u(r)$ must have the form $\exp (-r)$. So the radial part $R(r)$ is assumed to have the form

$$
\begin{equation*}
R(r)=\sum_{n}^{\pi} c_{n} r^{n} \exp \left(-\alpha r^{3 / 2}-\beta r\right) \tag{2.7}
\end{equation*}
$$

where $\alpha, \beta$ and $c_{n}$ are constants independent of $r$, and their values are dexermined to satisfy the Schrödinger equation (2.4). The constant $\alpha$ is of course positive, because the radial wavefunction $R(r)$ must become zero at $r \rightarrow \infty$. If the divergence of the radial nart $R(r)$ is prohibited at the origin $r=0, n$ must be larger than zero.

After substitution of (27) into (2.4), the left-hand side of (2.4) becomes
$\frac{1}{r^{2}} \frac{\mathrm{~d}}{\mathrm{~d} r}\left(r^{2} \frac{\mathrm{~d} R}{\mathrm{~d} r}\right)$

$$
\begin{align*}
= & \sum_{n} c_{n}\left\{\frac{9}{4} \alpha^{2} r^{n+1}+3 \alpha \beta r^{n+(1 / 2)}+\beta^{2} r^{n}-\frac{3}{2} \alpha\left(2 n+\frac{5}{2}\right) r^{n-(1 / 2)}\right. \\
& \left.-2 \beta(n+1) r^{n-1}+n(n+1) r^{n-2}\right\} \exp \left(-\alpha r^{3 / 2}-\beta r\right) \tag{2.8}
\end{align*}
$$

and the right-hand side of (24) is
$\sum_{n} c_{n}\left\{\frac{2 m}{\hbar^{2}} a r^{n+1}-\frac{2 m}{\hbar^{2}} E r^{n}-\frac{2 m}{\hbar^{2}} b r^{n-1}+l(l+1) r^{n-2}\right\} \exp \left(-\alpha r^{3 / 2}-\beta r\right)$.
Then the equation we have to solve is

$$
\begin{align*}
& \sum_{n} c_{n}\left[\left(\frac{9}{4} \alpha^{2}-\frac{2 m}{\hbar^{2}} a\right) r^{n+1}+3 \alpha \beta r^{n+(1 / 2)}+\left(\beta^{2}+\frac{2 m}{\hbar^{2}} E\right) r^{n}-\frac{3}{2}\left(2 n+\frac{5}{2}\right) \alpha r^{n-(1 / 2)}\right. \\
&-\left.\left\{2(n+1) \beta-\frac{2 m}{\hbar^{2}} b\right\} r^{n-1}+\{n(n+1)-l(l+1)\} r^{n-2}\right]=0 \tag{2.10}
\end{align*}
$$

This equation is rewritten into the following recurrence formula.

$$
\begin{align*}
& \{n(n+1)-l(l+1)\} c_{n}-\left(2 n \beta-\frac{2 m}{\hbar^{2}} b\right) c_{n-1}-\frac{3}{2}\left(2 n-\frac{1}{2}\right) \alpha c_{n-(3 / 2)} \\
& \quad+\left(\beta^{2}+\frac{2 m}{\hbar^{2}} E\right) c_{n-2}+3 \alpha \beta c_{n-(s / 2)}+\left(\frac{9}{4} \alpha^{2}-\frac{2 m}{\hbar^{2}} a\right) c_{n-3}=0 \tag{2.11}
\end{align*}
$$

## 3. Solution of the recarrence formula

If the radial wavefunction $R(r)$ is required to be normalizable, $c_{n}$ must be zero at a finite number of $n$. If all $c_{n}$ are zero for $n>k$, the recurrence formula (2.11) gives the following conditions:

$$
\begin{align*}
& \frac{9}{4} \alpha^{2}-\frac{2 m}{\hbar^{2}} a=0  \tag{31}\\
& 3 \alpha \beta=0  \tag{3.2}\\
& \beta^{2}+\frac{2 m}{\hbar^{2}} E=0  \tag{3.3}\\
& \frac{3}{2}\left(2 k+\frac{5}{2}\right) \alpha=0  \tag{3.4}\\
& 2(k+1) \beta-\frac{2 m}{\hbar^{2}} b=0 \tag{3.5}
\end{align*}
$$

and

$$
\begin{equation*}
k(k+1)-l(l+1)=0 \tag{3.6}
\end{equation*}
$$

where it is assumed that $c_{k} \neq 0$.
Equations (3.1) and (3.5) fix $\alpha$ and $\beta$, but these $\alpha$ and $\beta$ cannot satisfy (3.2) and (3.4). So we introduce additional counter terms in the potentiai $U(r)$, that is,

$$
\begin{equation*}
U(r)=a r-\frac{b}{r}+e r^{1 / 2}+f r^{-1 / 2} \tag{3.7}
\end{equation*}
$$

where constants $e$ and $f$ are determined to satisfy the Schrödinger equation (2.1). Then (3.2) and (3.4) are rewritten into

$$
\begin{equation*}
3 \alpha \beta-\frac{2 m}{n^{2}} e=0 \tag{3.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{3}{2}\left(2 k+\frac{5}{2}\right) \alpha+\frac{2 m}{\hbar^{2}} f=0 . \tag{3.9}
\end{equation*}
$$

Now the simultaneous equations we have to solve are (3.1), (3.3), (3.5), (36), (3.8) and (3.9). From (3.1) we obtain $\alpha$;

$$
\begin{equation*}
\alpha=\frac{2}{3} \sqrt{\frac{2 m}{\hbar^{2}} a} \tag{3.10}
\end{equation*}
$$

where $\alpha$ is positive. Also, from (35) we obtain $\beta$.

$$
\begin{equation*}
\beta=\frac{m}{(k+1) \hbar^{2}} b \tag{3.11}
\end{equation*}
$$

It seems as if $\beta$ was dependent on $k$. But $k$ is determined by (3.6) as

$$
\begin{equation*}
k=l \tag{312}
\end{equation*}
$$

Thus

$$
\begin{equation*}
\beta=\frac{m}{(l+1)^{n^{2}}} b \tag{313}
\end{equation*}
$$

is constant. The condition (312) requires that all $c_{n}$ should be zero for $n<k$ if we use it in (2.10).

The energy eigenvalue $E$ is given by (3.3);

$$
\begin{equation*}
E=-\frac{m}{2(l+1)^{2} \hbar^{2}} b^{2} \tag{314}
\end{equation*}
$$

which depends only or the angular momentum $l$. Additional constants $e$ and $f$ are determined by (3.8) and (3.9) as

$$
\begin{equation*}
e=\frac{b}{l+1} \sqrt{\frac{2 m}{\hbar^{2}} a} \tag{3.15}
\end{equation*}
$$

and

$$
\begin{equation*}
f=-\left(l+\frac{5}{4} ; \sqrt{\frac{2 \frac{2}{2}^{2}}{m}} a\right. \tag{3.16}
\end{equation*}
$$

These are dependent on $l$ In order that the Schrodinger equation with the continement potential given by (2.3) nas an analytical solution, the additional terms are necessary, as shown in (3.7), and these terms depend on the orbital angular momentum l.

## 4. Conclusion

It has been shown that the spherical symmetnc Schrödinger equation with the potential

$$
\begin{equation*}
U(r)=a r-\frac{b}{r}+\frac{1}{l+1} \sqrt{\frac{2 m}{\hbar^{2}}} a b r^{1 / 2}-\left(l+\frac{5}{4}\right) \sqrt{\frac{2 \hbar^{2}}{m} a r^{-1 / 2}} \tag{4.1}
\end{equation*}
$$

has an analytical solution. The radial wavefunction with orbital angular momentum / can be writen as

$$
\begin{equation*}
R(r)=c r^{2} \exp \left(-\alpha r^{3 / 2}-\beta_{i}\right) \tag{4.2}
\end{equation*}
$$

where

$$
\begin{equation*}
\alpha=\frac{2}{3} \sqrt{\frac{2 m}{\hbar^{2}} a} \tag{4.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\beta=\frac{m}{(l+1) \hbar^{2}} h . \tag{4.4}
\end{equation*}
$$

The constant $c_{l}$ is determuned by the normalization condition
The energy eigenvalue $E$ is given by

$$
\begin{equation*}
E=-\frac{m}{2(l+1)^{2} \hbar^{2}} b^{2} \tag{45}
\end{equation*}
$$

and depends only on $l$ The state is determined only by the orbital angular momentum $l$ and is degenerate $(2 l+1)$-fold.

In the case where the potential does not melude the Coulomb term proportional to $1 / r$, we may set $b=0$. From (35) or (4.4), $\beta=0$ Then, from (4.5), the energy eigenvalue is always zero This potential is the simplest, but is not very interesting. So we propose the asymptotically linear confinement potential (4.1). The most popular phenomenological potential that is used to calculate hadronic properties is the linear+ Coulomb type potential [4]. If necessary, we could add other terms, but the potential (4.1) is the simplest asympotically linear potential that we can solve analytically. Calculations of hadron properties will be carried out with this potential.

For the purpose of comparison, Eichten's potential [4]

$$
\begin{equation*}
U(r)=a r-\frac{b}{r} \tag{4.6}
\end{equation*}
$$

and the new proposed one (4i) are shown in figure 1 with the same values of the parameters $a$ and $b$. It will be easy to determine the vdlues of the parameters $a$ and $b$ to reproduce the physical quantities of hadrons.


Figure 1. Companson of the potentral (41) with $l=0(-)$ and Etchten's potental [4]


If the quark state is given by an analytical form, many physical properties of quarks can be easily calculated and understood. It must be worthwhile to investigate quarks in the potential given by (4.1)

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